

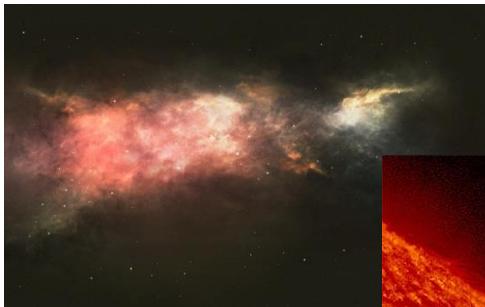
Alfvén wave parallel electric field in the dipole model of the magnetosphere

Kostarev D.V., Mager P.N., Klimushkin D.Yu.

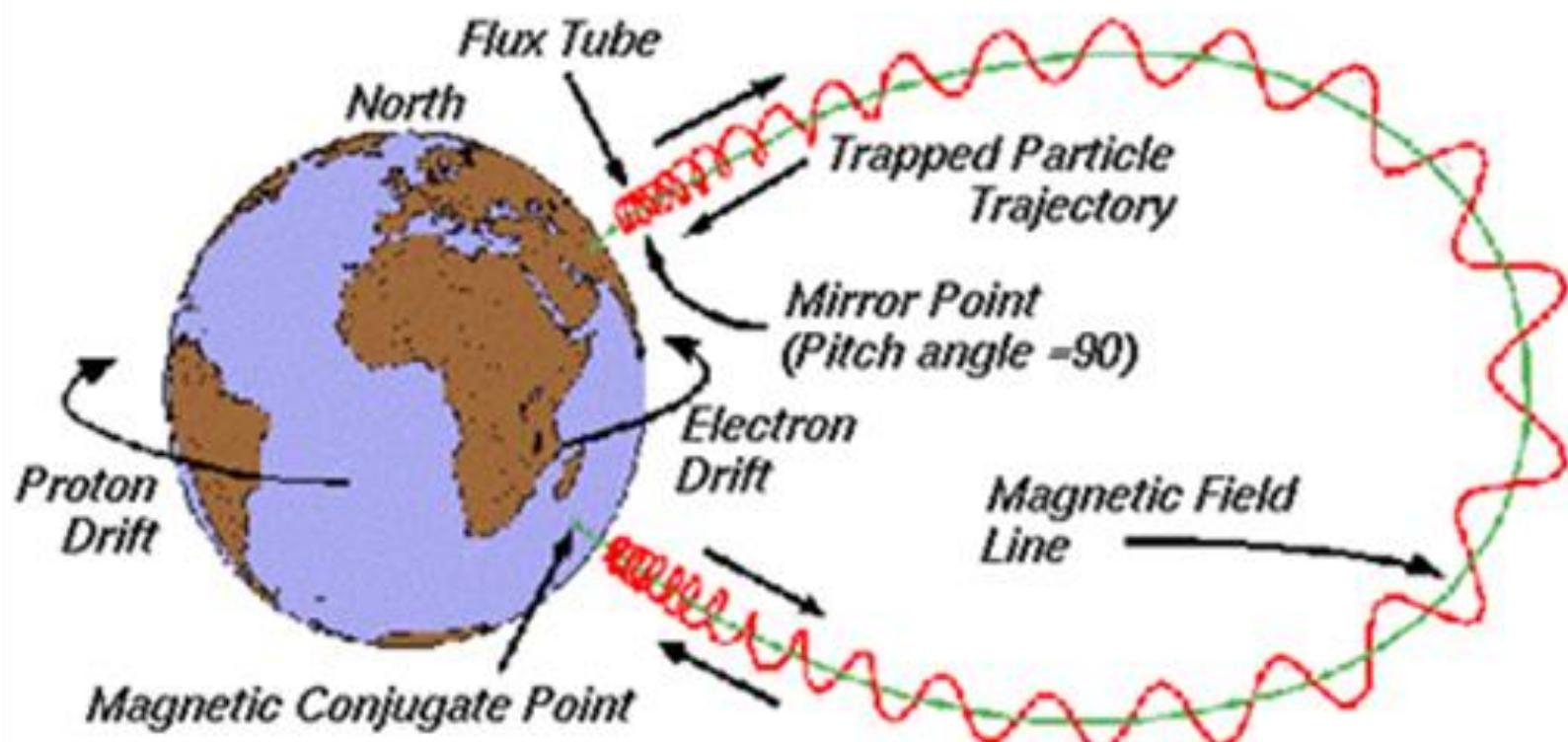
Institute of Solar-Terrestrial Physics of Siberian Branch of Russian Academy of Sciences, Irkutsk, Russia

ULF waves

(T > 0,2 sec)



Basic Components of Particle Motion: bounce, gyration and drift



Main steps

$$\left\{ \begin{array}{l} \hat{L}_A \psi + \hat{L}_C b_{\parallel} = 0 \\ \hat{L}_C \psi + \hat{L}_M b_{\parallel} = 0 \\ \hat{L}_{\varphi} \varphi_{\parallel} = \hat{L}_{\psi} \psi + \hat{L}_b b_{\parallel} \end{array} \right. \xrightarrow{\text{[Klimushkin and Mager, 2015]}} \left\{ \begin{array}{l} \psi \\ b_{\parallel} \approx \frac{c k_2}{\omega \sqrt{g_2}} \left(\frac{4\pi P'}{\sqrt{g_1} B^2} + \frac{3}{2} \frac{\beta}{1+\beta} \frac{1}{R} \right) \psi \\ \phi_{\parallel} \ll \psi \end{array} \right. \quad \boxed{\vec{E}_{\perp} = -\nabla_{\perp} \psi} \\ \boxed{\vec{E}_{\parallel} = -\nabla_{\parallel} \varphi_{\parallel}}$$

$$\sum_{i,e} \frac{q^2}{m} \int d\mu d\varepsilon \left(\frac{\partial F}{\partial \varepsilon} (\varphi_{\parallel} + \psi) - J_0^2 \frac{\hat{Q}F}{\omega} \psi + \frac{m}{q} J_0 \delta k_S \right) \frac{B}{|v_{\parallel}|} = 0$$

$$\delta k_{S_e} = -\frac{q_e}{m_e} \frac{\hat{Q}F_e}{\omega} \left(\varphi_{\parallel} + \frac{\omega_{d_e}}{\omega} \psi + \frac{m_e}{q_e} \mu b_{\parallel} \right) \quad \xrightarrow{\text{(...)} = \frac{2}{\tau_b} \int_{-l_0}^{l_0} \frac{(\dots)}{|v_{\parallel}|} dl_{\parallel}}$$

$$\hat{Q}F = \omega \frac{\partial F}{\partial \varepsilon} + \frac{k_2}{\omega_c \sqrt{g_{\perp}}} \frac{\partial F}{\partial x^1}$$

$$\tau_b = 2 \int_{-l_0}^{l_0} |v_{\parallel}|^{-1} dl_{\parallel}$$

Reference

Chen L., Hasegawa A. Kinetic theory of geomagnetic pulsations. 1. Internal excitations by energetic particles. *J. Geophys. Res.* 1991, vol. 96, pp. 1503–1512.
DOI:10.1029/90JA02346.

Klimushkin D. Y., and P. N. Mager The Alfvén mode gyrokinetic equation in finite-pressure magnetospheric plasma, *J. Geophys. Res. Space Physics* 2015, vol. 120, pp. 4465–4474. DOI:10.1002/2015JA021045.

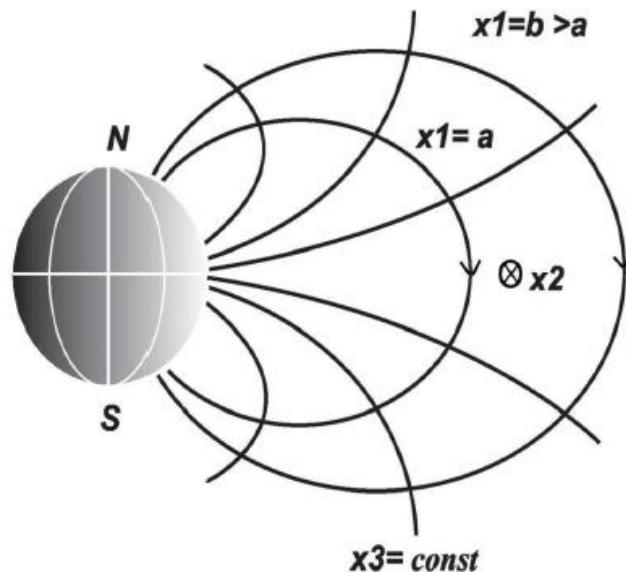
The distribution function

$$F = \frac{n}{(2\pi V)^{3/2}} \exp\left(-\frac{\nu^2}{2V^2}\right)$$

Approximation used for the solution

$$k_{\perp} \rho \ll 1 \quad \beta \ll 1 \quad \omega \gg \bar{\omega}_d$$

Coordinate system



Analytical solution

$$\phi_{\parallel} - \frac{1}{L} \int_0^{l_I} K(l, l') \phi_{\parallel}(l') dl' = \frac{\omega_{d_{e0}}}{\omega} \left[\frac{2\sqrt{g_1}}{R} \psi + \frac{3}{4L} \int_0^{l_I} \tilde{K}(l, l') \phi_{\parallel}(l') dl' \right]$$

$$K(l, l') = \theta(l - l') \frac{B(l)}{B_0} \int_{B_0/B_I}^{B_0/B(l)} \frac{d\lambda}{u(\lambda, l) u(\lambda, l')} + \theta(l' - l) \frac{B(l)}{B_0} \int_{B_0/B_I}^{B_0/B(l')} \frac{d\lambda}{u(\lambda, l) u(\lambda, l')}$$

$$\begin{aligned} \tilde{K}(l, l') &= \theta(l - l') \frac{B(l)}{B_0} 2 \frac{\sqrt{g_1}}{R} \left(\int_{B_0/B_I}^{B_0/B(l)} \frac{\lambda d\lambda}{u(\lambda, l) u(\lambda, l')} - \int_{B_0/B_I}^{B_0/B(l)} \frac{d\lambda}{u(\lambda, l) u(\lambda, l')} \right) + \\ &\quad \theta(l' - l) \frac{B(l)}{B_0} 2 \frac{\sqrt{g_1}}{R} \left(\int_{B_0/B_I}^{B_0/B(l')} \frac{\lambda d\lambda}{u(\lambda, l) u(\lambda, l')} - \int_{B_0/B_I}^{B_0/B(l')} \frac{d\lambda}{u(\lambda, l) u(\lambda, l')} \right) \end{aligned}$$

$$u(\lambda, l) = \sqrt{1 - \lambda B(l)/B_0}$$

Numerical solution

Poloidal wave case $k_1 = 0$

$$\vec{E}_a = 10 \text{ mV/m}$$

$k_2 = 50$ (*azimuthal wave number*)

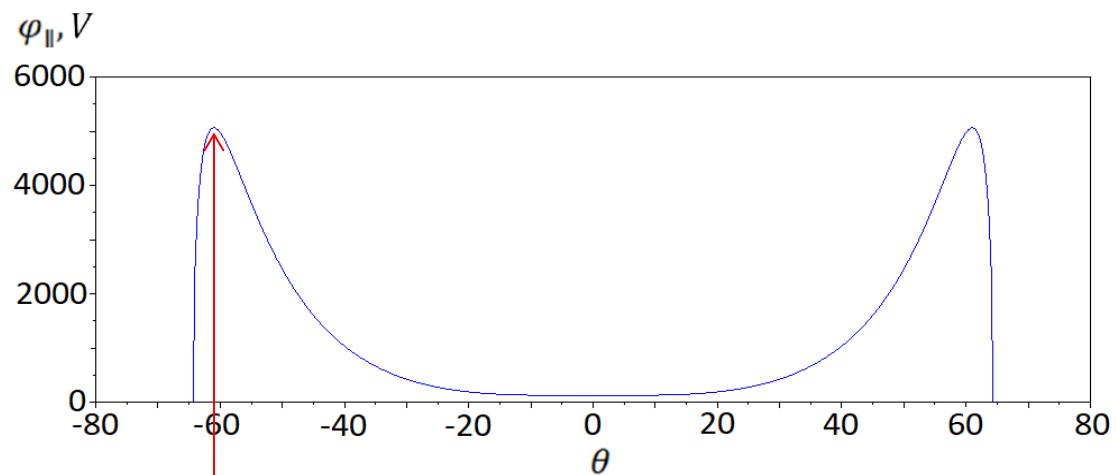
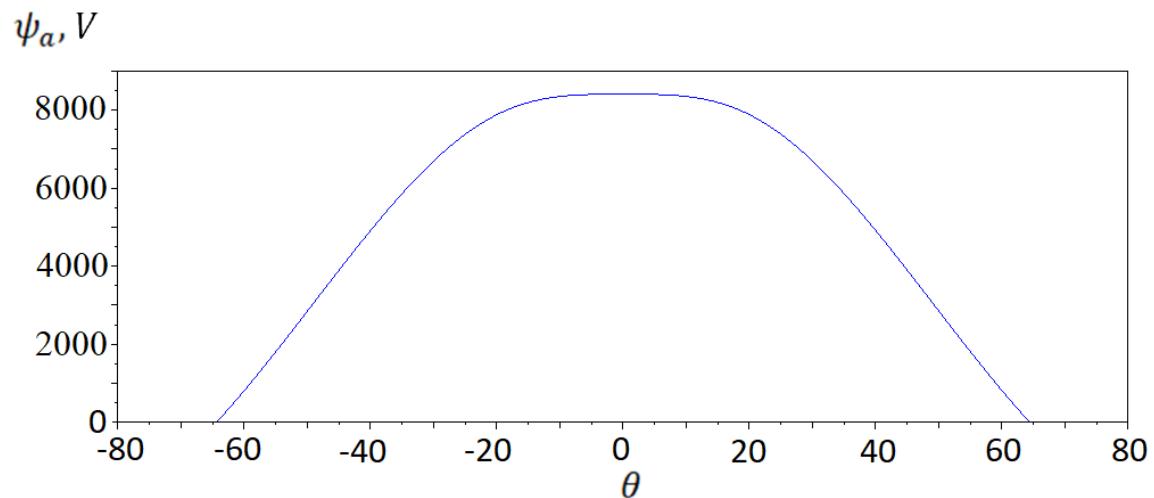
$$V_A = 1200 \text{ km/s}$$

$$L = 6.6 R_E$$

$$R_I = R_E + 1500 \text{ km}$$

$$\omega = 11.8 \text{ mHz}$$

$$T_e = 1 \text{ keV}$$



$\sim 0.5 R_E$

Thank you for your
attention!